# Application of the Method of Least Squares to the Elements of the Planet Pallas

1

In volume I of the *Gottingen Nachrichten*, Gauss gave the application of his method to the correction of the elements of the planet Pallas. Since the illustrious mathematician developed the algorithm outlined more briefly in his great work *Theoria Motus Corporum Coelestium* (see the previous Note) on this example, we felt we should translate here this part of his memoire. Since the first part requires an extensive knowledge of the theory of planetary motion, we shall not reproduce it, and we take as starting point the twelve equations which the six elements of the orbit ought to satisfy.

Denoting these corrections by

$$dL$$
,  $dZ$ ,  $d\pi$ ,  $dp$ ,  $d\Omega$ ,  $di$ 

the equations obtained by Gauss are the following

```
183.93''
               0.79363dL
                                  143.66dZ
                                                    0.39493d\pi
                                  0.18856d\Omega
                                                    0.17387di;
               0.95920d\phi
                                                +
6.81''
               0.02658dL
                                  46.71dZ
                                                    0.02658d\pi
                                                +
               0.20858d\phi
                                  0.15946d\Omega
                                                    1.25782di;
0.06''
               0.58880dL
                                  358.12dZ
                                                    0.26208d\pi
                              +
               0.85234d\phi
                                  0.14912d\Omega
                              +
                                                    0.17775di;
3.09''
               0.01318dL
                                  28.39dZ
                                                    0.01318d\pi
               0.07861d\phi
                                  0.91704d\Omega
                                                    0.54365di;
0.02''
               1.73436dL
                              +
                                  1846.17dZ
                                                    0.54603d\pi
               2.05662d\phi
                                  0.18853d\Omega
                                                    0.17445di;
8.98''
               0.12606dL
                                  227.42dZ
                                                +
                                                    0.12606 d\pi
               0.38939d\phi
                                  0.17176d\Omega
                                                    1.35441di;
2.31''
               0.99584dL
                                  1579.03dZ
                                                +
                                                    0.06456 d\pi
                              +
               1.99545d\phi
                                  0.06040d\Omega
                                                    0.33750di;
2.47''
               0.08089dL
                                  67.22dZ
                                                    0.08089d\pi
               0.09970d\phi
                                  0.46350d\Omega
                                                +
                                                    1.22803di;
0.01''
               0.65311dL
                                  1329.09dZ
                                                    0.38994d\pi
               0.08439d\phi
                                  0.04305d\Omega
                                                    0.34268di;
38.12''
               0.00218dL
                                  38.47dZ
                                                    0.00218d\pi
               0.18710d\phi
                                  0.47301d\Omega
                                                    1.14371di;
317.73''
               0.69957dL
                                  1719.32dZ
                                                +
                                                    0.12913d\pi
               1.38787d\phi
                              +
                                  0.17130d\Omega
                                                    0.08360di;
117.97''
               0.01315dL
                                  43.84dZ
                                                +
                                                    0.01315 d\pi
           +
               0.02929d\phi
                             +
                                  1.02138d\Omega
                                                    0.27187di.
```

From the nature of the observations which furnished the tenth of these equations, it is judged to inspire too little confidence to make use of it, and the six unknowns will be determined only from the other eleven.

The following explanations are literally translated from Gauss's Memoire

# 2 ( $\S 13$ in the original)

Since it is impossible for us to satisfy the eleven proposed equations exactly, that is to say, to make all the right hand sides zero, we shall seek to make the sum of their squares as small as possible.

One sees easily that if one considers the linear functions

$$n + ap + bq + cr + ds + \dots = w,$$
  
 $n' + a'p + b'q + c'r + d's + \dots = w',$   
 $n'' + a''p + b''p + c''r + d''s + \dots = w'',$ 

the equations which must be solved in order to make

$$\Omega = w + w' + w'' + \dots$$

a minimum, are

$$aw + a'w' + a''w'' + \dots = 0,$$
  
 $bw + b'w' + b''w'' + \dots = 0,$   
 $cw + c'w' + c''w'' + \dots = 0,$ 

or, defining the following abbreviations

$$an + a'n' + a''n'' + \dots = (an),$$
 $a^2 + a'^2 + a''^2 + \dots = (aa),$ 
 $ab + a'b' + a''b'' + \dots = (ab),$ 
 $\dots$ 
 $b^2 + b'^2 + b''^2 + \dots = (bb),$ 
 $bc + b'c' + b''c'' + \dots = (bc)$ 

p, d, r, s, etc. should be determined by the following equations

The process of solution, very tedious when the number of unknowns is considerable, can be simplified notably in the following way. Suppose that besides the coefficients (an), (aa), etc. (of which the number if  $\frac{1}{2}(i^2+3i)$ , if the number of unknowns is i) one has calculated the sum

$$n^2 + n'^2 + n''^2 + \dots = (nn);$$

one sees easily that one has

$$\Omega = (nn) + 2(an)p + 2(bn)q + 2(cn)r + \dots + (aa)p^2 + 2(ab)pq + 2(ac)pr + \dots + (bb)q^2 + 2(bc)qr + 2(bd)qs + \dots + (cc)r^2 + 2(cd)rs + \dots;$$

and, denoting

$$(an) + (aa)p + (ap)q + \dots$$

by A, all the terms of  $\frac{A^2}{(aa)}$  which contain the factor p, are found in the expression  $\Omega$ , and consequently

 $\Omega - \frac{A^2}{(aa)}$ 

is a function independent of p. This is why, setting

one has

$$-\frac{A^{2}}{(aa)} = (aa,1) + 2(bn,1)q + 2(cn,1)r + 2(dn,1)s \dots + (bb,1)q^{2} + 2(bc,1)qr + 2(bd,1)qs \dots + (cc,1)r^{2} + 2(cd,1)rs \dots + \dots$$

We shall denote this function by  $\Omega'$ .

Similarly, setting

$$(bn,1) + (bb,1)q + (bc,1)r + \dots = B,$$

the difference

$$\Omega' - \frac{B^2}{(bb,1)}$$

will be independent of q; we shall represent it by  $\Omega''$ .

Similarly, setting

$$\begin{array}{rclrcl} (nn,1) & - & \frac{(bn,1)^2}{(bb,1)} & = & (nn,2), \\ (cn,1) & - & \frac{(bn,1)(bc,1)}{(bb,1)} & = & (cn,2), \\ (cc,1) & - & \frac{(bc,1)^2}{(bb,2)} & = & (cc,2), \\ & & & & & & & \\ (cn,2) & + & (cc,2)r & + & (cd,2)s & + & \dots & = & C \\ \end{array}$$

the difference

$$-\frac{C^2}{(cc,2)}$$

will be a function independent of r.

Continuing in this way we shall form a sequence of expressions, ', ", etc., of which the last will be independent of the various unknowns and is denoted by (nn, ), if denotes the number of these unknowns; then we shall have

$$\Omega = \frac{A^2}{(aa)} + \frac{B^2}{(bb,1)} + \frac{C^2}{(cc,2)} + \frac{D^2}{(dd,3)} + \dots + (nn,\mu)$$

One can easily prove that since is a sum of squares

$$w^2 + w'^2 + w''^2 + \dots$$

and cannot become negative, the denominators (aa), (bb,1), (cc,2), etc. are all positive. (For brevity we omit the details of the proof.) Accordingly, the minimum value of obviously corresponds to the values of the unknowns for which

$$A = 0, \quad B = 0, \quad C = 0, \quad ...,$$

and, by starting to solve the system with the last equation, which contains only one of them, one finds the values of p, q, r, s, etc. without having to carry out any elimination. The method gives at the same time the minimum value of , which is (nn, ).

# 3 ( $\S$ 14 in the original)

Let us apply these principles to our example, in which p, q, r, s, etc. are replaced by  $dL, dZ, d\pi, d\phi, d\Omega, di$ . By careful calculation I have found

```
(nn)
          +148848
                     (ac)
                               -0.09344
                                            (cc) =
                                                    +0.71917
(an)
          -371.09
                               -2.28516
                                            (cd) =
                                                    +1.13382
                     (ad)
(bn)
          -580104
                               -0.34664
                                            (ce) =
                                                    +0.06400
                     (ae)
          -113.45
                               -0.18194
                                            (cf) =
(cn)
                     (af)
                                                    +0.26341
(dn)
          +268.53
                     (bb)
                               +10834225
                                            (dd) =
                                                    +12.00340
(en)
          +94.26
                     (bc)
                               -49.06
                                            (de) =
                                                    -0.37137
(fn)
          -31.81
                     (bd)
                               -3229.77
                                            (df) =
                                                    -0.11762
                                            (ee) =
(aa)
         +5.91569
                     (be)
                               -198.64
                                                    +2.28215
(ab)
          +7203.91
                     (bf)
                               -143.05
                                            (ef) = -0.36136
(ff)
          -5.62456
```

from which one obtains

### Similarly

#### From which;

### Similarly:

$$(nn,4) = +98963$$
  $(fn,4) = +4.33$   $(ef,4) = -0.37766$   $(en,5) = +75.23$   $(ee,4) = +2.22346$   $(ff,4) = +5.42383$ 

### From which:

$$(nn,5) = +96418 \quad (fn,5) = +17.11 \quad (ff,5) = +5.42383$$

### From which finally we obtain

$$(nn, 6) = +96364$$

## Thus we have the following six equations

```
= +17.11''
                     +5.42383di
0
   =
       +75.23''
                     +2.22346d\Omega
                                    -0.37766di
0
       +25.66''
                     +9.29213d\phi
                                    -0.36175d\Omega
                                                   -0.57384di
0
       -115.81''
                     +0.71612d\pi
                                    +1.11063d\phi
                                                   -0.06392d\Omega +0.25868di
0
      -13854''
                     +2458225dZ
                                    +62.13d\pi
                                                   -0.51058 d\phi
                                                                 +213.84d\Omega
       +73.45 di
       -371.09''
                     +5.91569dL
                                    +7203.91dZ -0.00344d\pi -2.20516d\phi
       -0.34664d\Omega -0.18194di
```

#### from which one obtains;

$$di = -3.15''$$

$$d\Omega = -34.37''$$

$$d\phi = -4.29''$$

$$d\pi = +166.44''$$

$$dZ = +0.054335''$$

$$dL = -3.06''$$

These are the corrections which should be applied to the elements originally found for the planet.

Taken from *Work (1803–1826) on the Theory of Least Squares*, trans. H F Trotter, Technical Report No.5, Statistical Techniques Research Group, Princeton, NJ: Princeton University 1957.