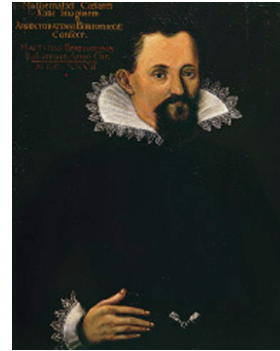




# Lecture 7. Kepler: Laws of Planetary Motion

## Johannes Kepler (1571-1630)



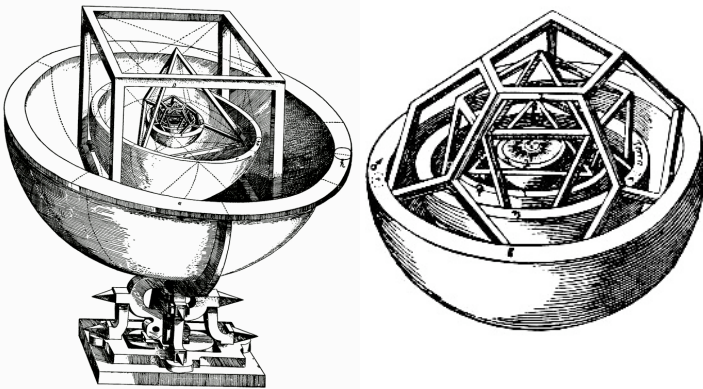
### Kepler

- Johannes Kepler came from a poor Protestant family in Germany.
- He became aware of Copernicus' work at the University of Tübingen, where he completed a master's degree.
- He then began studying theology, but before finishing his degree he was offered a job teaching mathematics at a (Lutheran) school in Austria.
- Kepler was impressed with the Pythagorean idea that numbers underlie the mysteries of the cosmos.
- He was especially interested in why there are only 5 planets, and why they orbit at the distances they do from the Sun.

### *Mysterium Cosmographicum*

- While teaching astrology, Kepler was struck with a great idea.
- There are 5 regular solids: the cube, the tetrahedron, the octahedron, the dodecahedron, and the icosahedron.
- If the orbit of each planet were circumscribed by one of these regular solids, and each of these was nested inside the next, he achieved a solar system with distances close to the distances calculated by Copernicus accurate to about 5%.
- This then accounted for the number of planets as well as their distances from the Sun.
- "I wanted to become a theologian, and for long I was uneasy, but now, see how through my efforts God is being celebrated in astronomy."

### Kepler's solar system



### Kepler

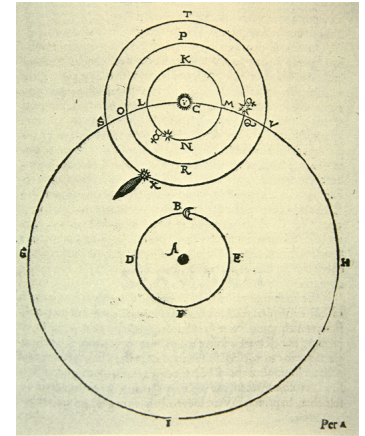
- Published *Mysterium cosmographicum* in 1597.
- N.B. Kepler is doing something that marks the beginning of a more empirical approach to cosmology: he is testing his model against measurements.
- The book was read and added support to the Copernican movement.
- In 1598 Kepler was forced to leave Graz when he refused to convert to Catholicism.
- In 1600 he finally got a job with the great observational astronomer Tycho Brahe in Prague at the court of Rudolf II.
- His job was to use Tycho's data on Mars to fit with Tycho's cosmological model.

## Tycho Brahe (1546-1601)



## Tycho's cosmos

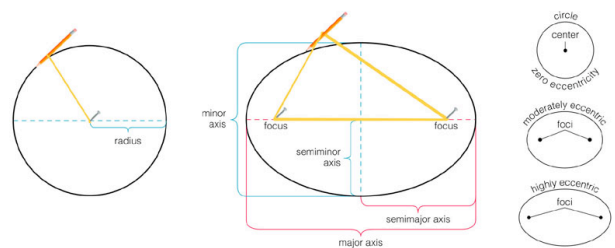
- Tycho had made the most accurate observations obtained at that time on the planets.
- Tycho's model of the heavens had the planets orbiting round the Sun and the Sun orbiting round the Earth at the center of the Universe.



## Kepler's laws of planetary motion

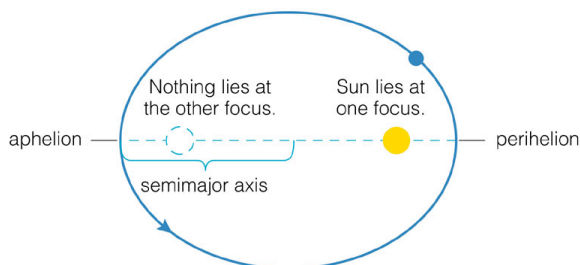
- Tycho died after two years and Kepler inherited his data and his title: Imperial Mathematician to the Emperor Rudolf II.
- Kepler dutifully attempted to reconcile the Mars data using models of Ptolemy, Copernicus, and Tycho.
- None were successful at representing Tycho's accurate data for Mars.
- After six years of work he gave up attempting to use circles for the planetary orbits.
- Kepler realized Mars moves in an ellipse around the Sun.
- In his *Astronomia Nova* (1609) he presented his first two laws of planetary motion.

## Geometry of ellipses



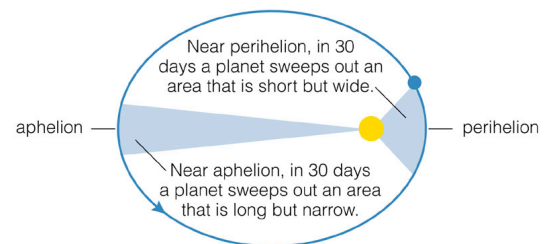
## Kepler's First Law

First Law: Planets move in **elliptical orbits with the Sun at one focus of the ellipse.**



## Kepler's Second Law

Second Law: A line from the Sun to the planet sweeps out **equal areas in equal times**, i.e. planets don't move at constant speed.



The areas swept out in 30-day periods are all equal.

## Kepler's Third Law

- Kepler was a committed Pythagorean, and he searched for 10 more years to find a mathematical law to describe the motion of planets around the Sun.
- In *Harmony of the World* (1619) he enunciated his Third Law:
- (Period of orbit)<sup>2</sup> proportional to (semi-major axis of orbit)<sup>3</sup>.
- In symbolic form:  $P^2 \propto a^3$ .
- If two quantities are proportional, we can insert a proportionality constant,  $k$ , which depends on the units adopted for  $P$  and  $a$ , and get an equation:
  - $P^2 = ka^3$ .

## Using Kepler's 3rd Law

Kepler's third law:  $P^2 \propto a^3$

Therefore  $P^2 = ka^3$ , where  $k$  is some constant number.

We can find  $k$  if  $P$  is expressed in (Earth's orbit) years and  $a$  is expressed in terms of the distance between Sun & Earth. This distance is called 1 Astronomical Unit, or 1 A.U.

Then substituting above values into  $P^2 = ka^3$  we find:

$$(1)^2 = k(1)^3.$$

Solving for  $k$ , we find  $k = 1$ .

When these units for  $P$  and  $a$  are used:

$$P^2 = a^3 \text{ with } P \text{ in years and } a \text{ in A.U.}$$

## Example

An asteroid is found and its orbital semi-major axis around the Sun is measured to be 4 A.U. What is the period of its orbit round the Sun?

$$P^2 = a^3 \text{ with } P \text{ in years and } a \text{ in A.U.}$$

$$\text{Since } a = 4 \text{ A.U.},$$

$$P^2 = a^3 = (4)^3 = 64.$$

$$\text{So } P = 8 \text{ years.}$$

## Celestial dynamics?

- But what makes the planets move in this way?
- Kepler has now lost Aristotle's theory of motion, and there was nothing to replace it.
- He realized that planets move more slowly the further they were away from the Sun.
- Does this mean it is the Sun that is the source of their motion?
- Kepler guessed that the planets are moved in some way by the Sun, and the further away a planet is, the slower it orbits.
- He thought it might be some kind of magnetic force.

## Kepler's intellectual legacy

- The three laws of planetary motion are an intellectual high water mark.
- The Sun was now firmly at the center of the Solar system.
- For the first time, Kepler introduced celestial motion that was **not circular**. (The Greek idea that motion in the heavens must be circular had lasted over 2,000 years.)
- Kepler was aware of the fact that his model was lacking in dynamics. He had discarded Aristotle's model, and with it his theory of motion, but **there was now no theory to say what made the planets move in these ellipses**.
- Kepler suggested there was some magnetic force emanating from the Sun.
- This became **a major intellectual problem in the 17th C.**