JAXA Report on Hayabusa-2, Procyon, and International Collaboration Sample Return Working Group

May 21st, 2014

NASA-HQ (Rm 3D42) via Teleconference from JAXA/ISAS

Hajime YANO (JAXA/ISAS, Japan)

About the Presenter



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JAXA Space Exploration Center (JSPEC)

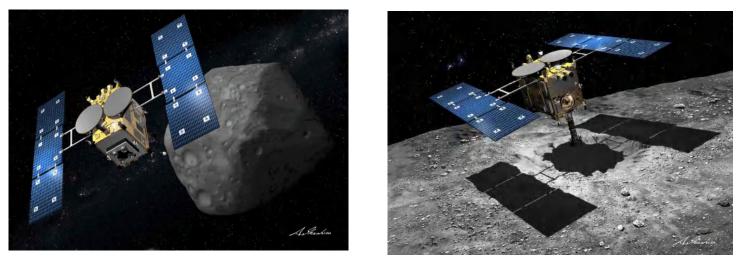
<Projects, WGs, Committee>

- Hayabusa Project Team (1996-2013)
- IKAROS Project Team (2007-2013)
- Enceladus Sample Return Feasibility Research Team (2011-2013)
- •Tanpopo Project Team (2007~)
- Hayabusa-2 Project Team (2011~)
- Solar Power Sail Working Group (2002~)
- International Collaboration Sample Return Working Group (2013~)
- JAXA Planetary Protection Safety Review Board (2013~)



NASA Advisory Council Planetary Protection Subcommittee May 2014

Hayabusa-2 Probability of Impact Analysis for Planetary Protection (Status Report)

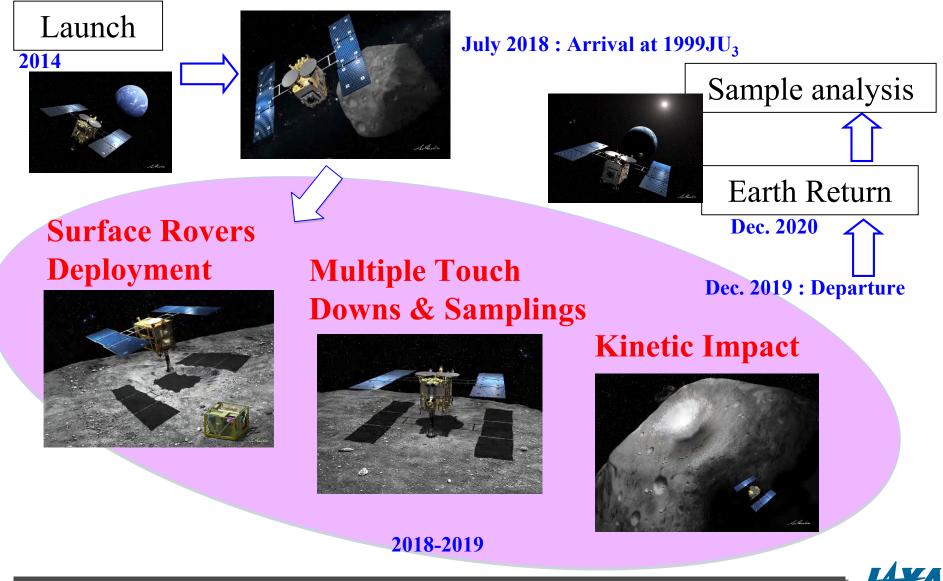


May 21st, 2014 NASA-HQ (Rm 3D42) via Teleconference from JAXA/ISAS

Presented by Hajime YANO (JAXA/ISAS, Japan) On behalf of the Hayabusa-2 Project Team (Contributed by Y. Tsuda, T. Chujo and H. Yano)



Hayabusa-2 Mission Outline





Hayabusa-2 Mission and Trajectory

◆Hayabusa2 is the 2nd Japanese sample return mission to small body. JAXA will launch Hayabusa2 in 2014, explore C-type asteroid 1999JU3, and return back to the Earth in 2020.

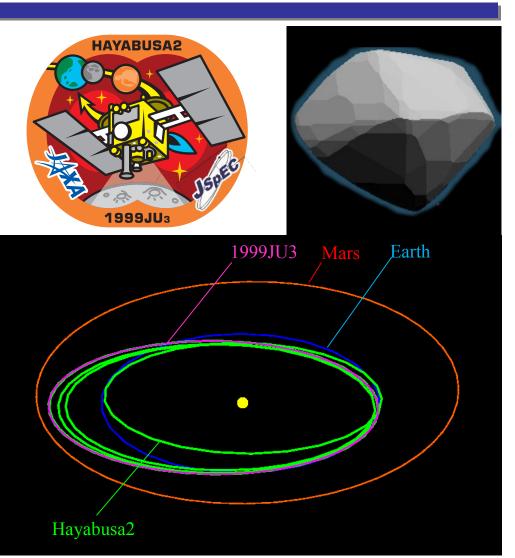
• Orbit of Asteroid 1999JU ₃ :	cf. Mars
Perihelion 0.96AU	$r_a = 0.093 AU$
Aphelion 1.42AU	$r_p = 1.67 AU$
Eccentricity 0.19	inc=1.85deg
Inclination 5.88deg (w.r.t. E	Celiptic plane)

*1999JU₃ itself is far enough (i.e. has zero collision probability) both from the Earth & Mars.

♦ Mission Schedule

Earth Departure:	Nov-Dec, 2014
Earth Swing-by:	Dec, 2015
1999JU ₃ Arrival:	Jul, 2018
1999JU ₃ Dep.:	Dec, 2019
Earth Reentry:	Dec, 2020

*Two backup windows are planned (May-Jun 2015 and Nov-Dec 2015), for which PP analyses are also conducted.



The entire mission trajectory



Planetary Protection Activity of Hayabusa-2

- ✓ In 2012, the Hayabusa-2 project requested the COSPAR Planetary Protection Panel to recommend both outbound and inbound planetary protection requirements to the 1999 JU3 sample return mission.
- ✓ After careful consideration with respect to the COSPAR Planetary Protection Policy at the Alpbach Colloquium in 2012 followed by the COSPAR2012 GA in India, <u>the</u> <u>COSPAR PPP concluded that the Hayabusa-2 mission was considered the outbound</u> <u>Category-II, in condition of "avoiding impact with Mars under all mission</u> <u>scenarios", and the inbound Category-V(B) as "unrestricted Earth return" for the</u> <u>1999 JU3 samples.</u>
- ✓ The above agreement also <u>concurred with the recommendation of the NASA</u> <u>Advisory Council's Planetary Protection Subcommittee in 2012.</u>
- ✓ The objective of today's report is to explain the last outstanding condition for completing the COSPAR requirements to the Hayabusa-2 mission:
- → To prove that its impact probability analysis to Mars in all the mission scenarios is less than 1% in 50 years since its launch.



Probability Analysis – Methodology

- Challenges in analyzing the Mars impact probability for Hayabusa-2
 - The trajectory is controlled mainly by low thrust propulsion (Ion Engine System: IES).
 - \rightarrow Non ballistic trajectory makes the 50yr probability analysis difficult.
 - The trajectory plan includes the Earth swing-by and Earth Return.
 - \rightarrow Probability of the Mars impact after the Earth gravity assist may not be negligible.
- Analysis Methodology: Piecewise Impulsive Approximation
 - <u>The IES continuous thrust history is partitioned to small nodes</u>, and the thrust in each node is represented by the equivalent impulsive ΔV at the center of the nodes.
 - The each impulsive ΔV is then treated as "TCM" of the ordinary trajectory deviation analysis for impulsive ΔV trajectory.
 - <u>The Mars impact after Earth swing-by case is evaluated independently of the direct Mars impact case</u>. An upper bound of the probability is evaluated employing an analytical/numerical-combined method. *(Result not shown in this preliminary report)*
- Analysis Strategy: Three-Step Analysis to reduce the computational burden.
- 1) Phase-Free Analysis (Mars true anomaly is neglected)
- 2) High Fidelity Monte-Carlo Simulation (Validation of step-1 for typical case only)
- 3) Mars Impact via Earth Swing-by Evaluation

$$P_{\rm I} = \int (p_{\rm drct} + p_{\rm swby}) q dt$$

 P_{I} : Overall Mars impact probability (for 50yrs), q(t): Spacecraft failure rate

JA XA

Mars Impact Probability (Preliminary)

The result at present considers "Direct Impact" case only. The Earth-swingby case is to be added. However, it is expected that the final result well satisfies $P_{\rm I}$ <<1e-2.

Trajectory Leg	$\begin{array}{c} \mathrm{IES} \ \Delta \mathrm{V} \ / \ \mathrm{RCS} \\ \Delta \mathrm{V} \end{array}$	Impact Probability (direct) $\int p^{dret} dt$	Impact Probability (Via Earth Swby) ∫ p ^{swby} idt	Reliability Based qi+1*a	Reliability based $\int (p^{drct} + p^{swby}) dt$
Injection	No / No	1.0e-14	To be calculated	1.0e-5	1.0e-19
Earth to Earth	Yes / Yes*c	4.0e-13	To be calculated	3.0e-2	1.2e-14
Earth to Asteroid	Yes / Yes*c	0* ^b	To be calculated	1.1e-1	0* ^b
Asteroid Proximity	No / No	0* ^b	To be calculated	4.1e-2	0* ^b
Asteroid to Earth	Yes / Yes*c	0* ^b	To be calculated	4.2e-2	0* ^b
TOTAL					1.2e-14

Table [N]: Nominal Window (Launch: Nov-Dec. 2014)

*a (Very conservative) upper bound

*b Negligible (<1e-10)

*c Very small RCS TCMs exist which are negligible compared with IES ΔV

Table [BU1]: Backup Window1 (Launch: May-Jun. 2015) *TOTAL Only

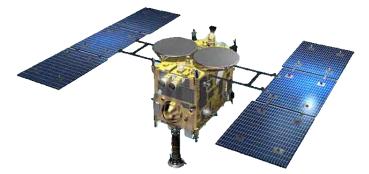
Trajectory Leg	$\begin{array}{c} \mathrm{IES} \ \Delta \mathrm{V} \ / \ \mathrm{RCS} \\ \Delta \mathrm{V} \end{array}$	Impact Probability (direct) $\int p^{drct} dt$	Impact Probability (Via Earth Swby)*a ∫p ^{swby} idt	Reliability Based qi+1	Reliability based $\int (p^{drct}_i + p^{swby}_i)qdt$
TOTAL					1.2e-14

Table [BU2]: Nominal Window (Launch: Nov-Dec. 2015) *TOTAL Only

Trajectory Leg	$\begin{array}{c} \mathrm{IES} \ \Delta \mathrm{V} \ / \ \mathrm{RCS} \\ \Delta \mathrm{V} \end{array}$	Impact Probability (direct) ∫ p ^{dret} idt	Impact Probability (Via Earth Swby)* ^a ∫ p ^{swby} idt	Reliability Based qi+1	Reliability based $\int (p^{drct} + p^{swby}) dt$
TOTAL					1.2e-14

NASA Advisory Council Planetary Protection Subcommittee May 2014

Probability Analysis for an Accidental Impact on Mars by PROCYON, Hayabusa-2's Piggyback Micro-spacecraft (Status Report)





May 21st, 2014

NASA-HQ (Rm 3D42) via Teleconference from JAXA/ISAS

Presented by Hajime YANO (JAXA/ISAS, Japan) On behalf of the PROCYON Project Team

(Contributed by R. Funase, Y. Kawakatsu, N. Ozaki, K. Ariu, S. Nakajima, Y. Shimizu, H. Yano)

Overview of the PROCYON Mission

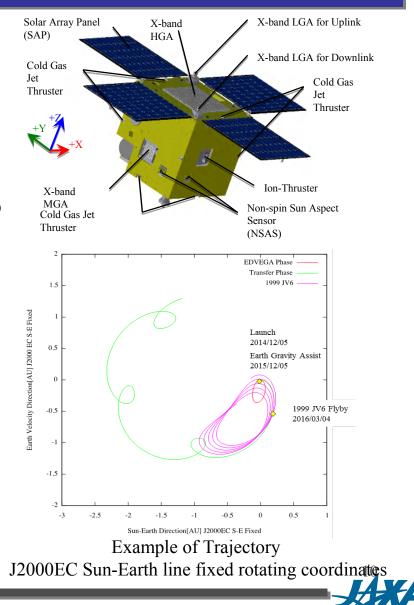
Mission Scenario
[Nominal] 1-year Electric Delta-VEGA (EDVEGA)
- Launch: $2014/12 \rightarrow$ Earth Gravity Assist (EGA) : $2015/12$
\rightarrow NEAs Flyby: 2016/TBD
[Backup 1] Direct Transfer
- Launch: $2014/12 \rightarrow NEAs$ Flyby: $2015/TBD$
[Backup 2] 2-year EDVEGA
- Launch: $2014/12 \rightarrow EGA$: $2016/12 \rightarrow NEAs$ Flyby: $2017/TBD$

Mass	Initial Total Mass[kg]: 67.0 Fuel Mass(Xe)[kg]: 2.5
Ion Propulsion System (IPS) using Xe	Thrust[mN]: 0.25 Isp[s]: 1000.0
Reaction Control System (RCS) using Xe	Thrust[mN]: 19.0 Isp[s]: 24.2

*<u>IES maximum $\Delta V = 235[m/s]$ as 2-years (mission life time)</u> <u>continuous maneuver.</u>

The phase-free minimum ΔV for Direct Impact on Mars is <u>250[m/s]</u> as impulsive maneuver.

→ PROCYON does not have a capability to direct impact on Mars

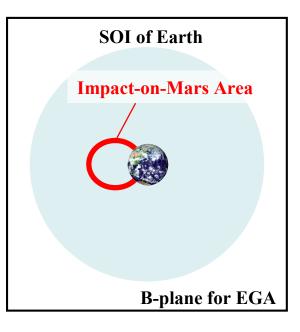


Probability Analysis for Impact on Mars via EGA — Phase-Free and Ballistic Analysis —

- Analysis Methodology
 - Evaluate B-plane position of EGA when the spacecraft crosses Mars SOI × 2,
 - With considering ballistic after EGA
 - Without considering the Mars position in the orbit (Phase-free)
- Initial Condition
 - Date-time of EGA: 2015-Dec-05 00:00:00
 - Position and velocity of the Earth is given by JPL ephemeris (DE405).
 - **V**∞ = [1.16, 2.59, -3.59] [km/s] @ J2000 EC
 - $||V\infty|| = 4.58[km/s]$

We assume that the probability of impact on Mars via EGA as follows:

 $\frac{\text{Probablity of Impact} = \frac{\text{Impact on Mars Area}}{\text{Area of Earth SOI}}$



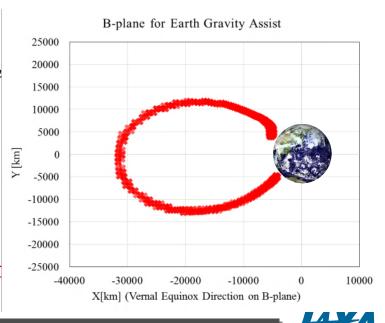
Analytical Results and Conclusion



depending on launch date/ launch error

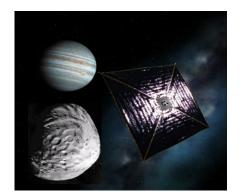
Thruster: MIPSThruster: MIPS (almost Ballistic)Thruster: CGJOrbit Determination: R&RROrbit Determination: R&RROrbit Determination: Optical Image

- Results
 - In the phase-free and ballistic analysis, the <u>probability of impact</u> on Mars via EGA is about 3.0×10^{-5} , which is by far less than 10^{-2}
- Conclusion
 - We conclude that the possibility of Mars impact is negligible within 50 year after its launch.
 - As a next step, we will analyze that the probability of <u>impact</u> <u>considering TCM guidance error before EGA</u>
 - Depending on the situation, we will eliminate the candidate
 NEAs if the spacecraft has too large probability of impact on Mar in the trajectory



NASA Advisory Council Planetary Protection Subcommittee May 2014

JAXA/ISAS International Collaboration Sample Return Working Group for Astrobiology-Driven Explorations



NASA-HQ (Rm 3D42) via Teleconference from JAXA/ISAS

May 21st, 2014

Presented by Hajime YANO (JAXA/ISAS, Japan)

On behalf of the JAXA/ISAS International Collaboration Sample Return WG (Contributed by H. Kuninaka, H. Yano, Y. Shimizu, T. Yamada, K. Yamada, M. Ohtsuki, H. Sawada, K. Takai, Y. Takano, T. Shibuya, Y. Sekine, R. Funase, etc.)

Japan's Past Solar System Missions Sorted by Destinations

Moon

Planets





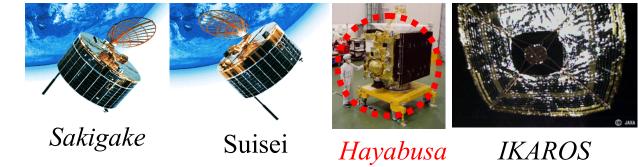
Hiten/Hagoromo

(LUNAR-A)



Kaguya



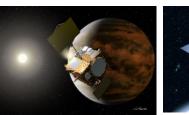


* Technology verification missions in Italic. International collaboration in red. PP activities in circle.

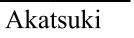
Japan's Current and Approved Missions Sorted by Destinations

Moon

Planets



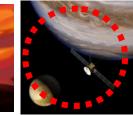




Hisaki

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Small Bodies

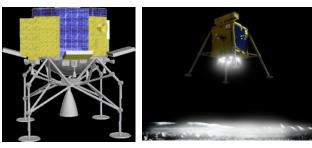


* Technology verification missions in Italic. International collaboration in red. PP activities in circle.



Japan's Proposed Future Plans Sorted by Destinations

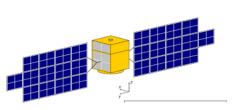
Moon



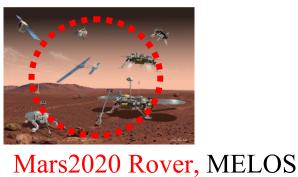
SLIM

SELENE-2

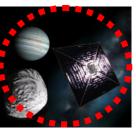
Planets



DESTINY



Small Bodies



Solar Power Sail

* Technology verification missions in Italic. International collaboration in red. PP activities in circle.



Future Direction-1: Deep Space Round Trips: Small Body Sample Returns by Hayabusa Heritage

Hayabusa Itokawa = S type (2003-10)

Post Hayabusa Era

Hayabusa-2

1999 JU3 = C type Lessons Learned from Hayabusa (2014-20)

D type

And Beyond

Solar Power Sail D type Jupiter Trojan Hayabusa + IKAROS **Technology (Early 2020's)**



Carbonaceous

(2016-23)

Chondrites

Ordinary Chondrites

S type

C type

Main Asteroid Belt



CP-IDP, Ultracarbonaceous AMMs? **Future Targets: Comet Nucleus, Icy Plymes**

Future Direction-2: Going to Outer Planet Regions: Building up TRLs of Solar Sail Technology



2003. August Balloon Test(B30-71) at 36km alt.: Active Deployment of Sail (4m)



2004. August Sounding Rocket(S310-34) at >100km alt.: Active Deployment of Sail (10m) Modeling of Sail Dynamics



M-V-7 Rocket Sub-payload (SSSAT) in LEO: Deployment Demo of Small Power Sail (5m)

2010. May

H-IIA-17 Piggy-back (IKAROS) in deep space: First Solar Sail in Interplanetary Space Deployment of Sail Membrane (200 m^2)

Early 2020's

Solar Power Sail (3000 m²) with Ion Engines: Cruising Science (IR astronomy, High energy astrophysics, Dust) and Jupiter and Trojan explorations

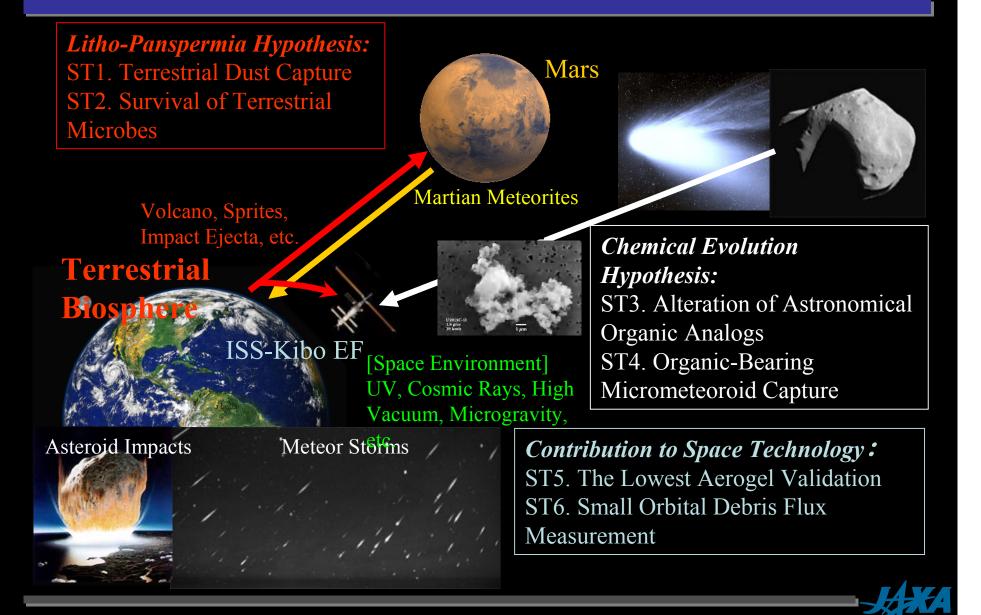




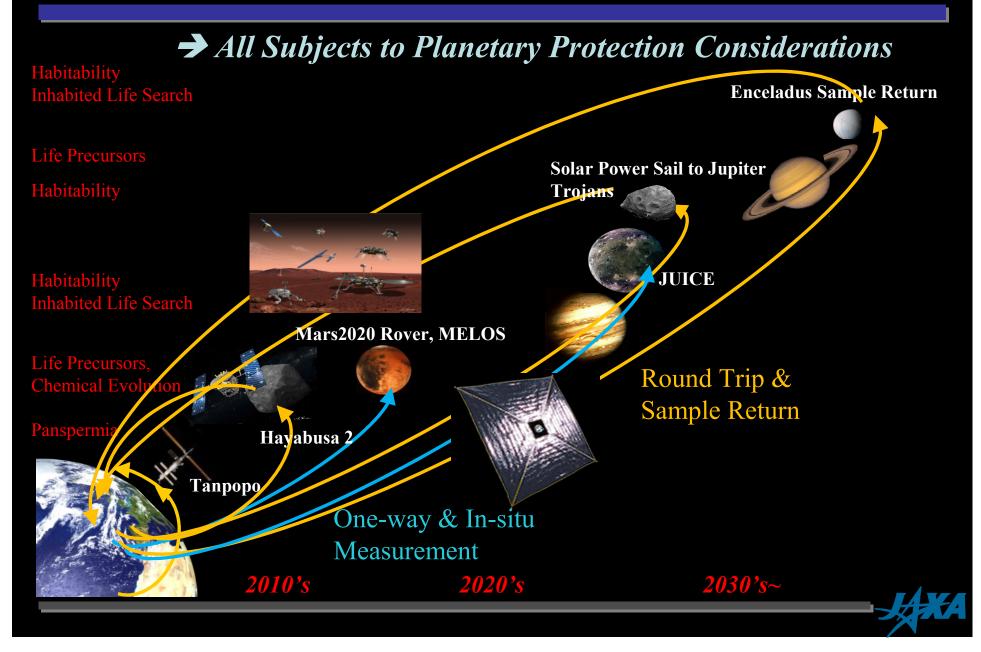




Future Direction-3: Astrobiology Driven Missions: Tanpopo: Astrobiology Exposure & Meteoroid Capture onboard ISS

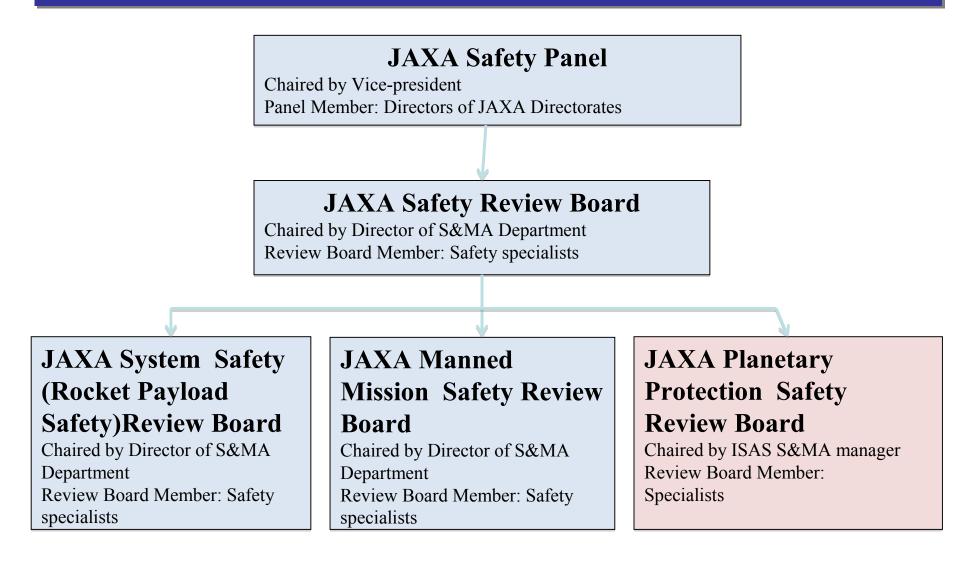


Solar System Exploration Vision in Japan: Keywords: Round Trip, Outer Planets, Astrobiology and International



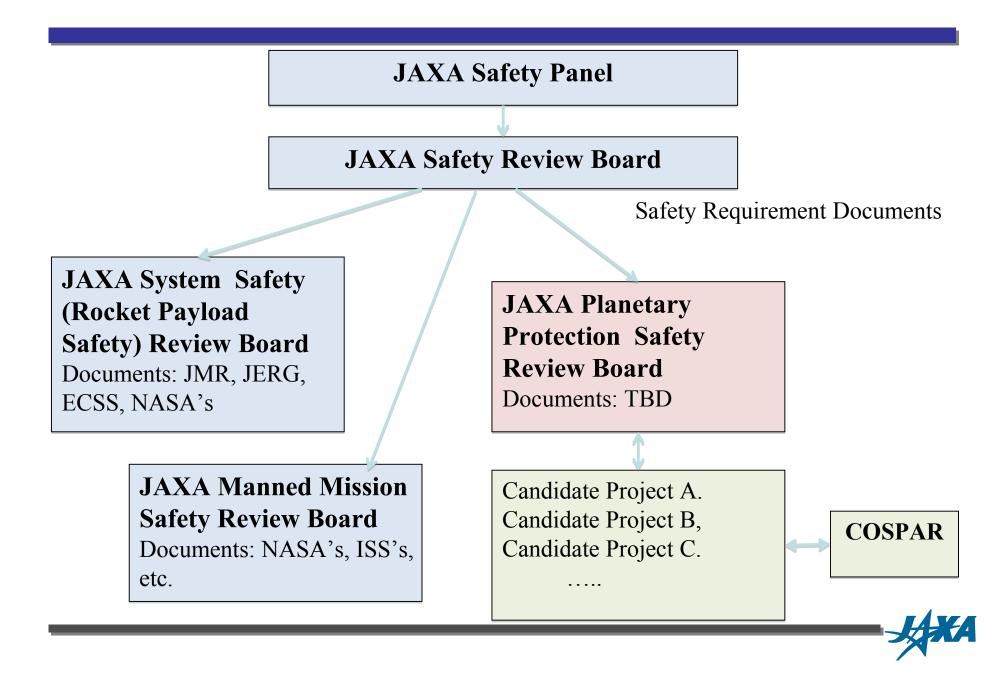
JAXA Safety Review Board Structure:

Planetary Protection Safety Review Board Established in 2013





JAXA Planetary Protection Safety Review Process



ISAS International Collaboration Sample Return WG Established in 2013

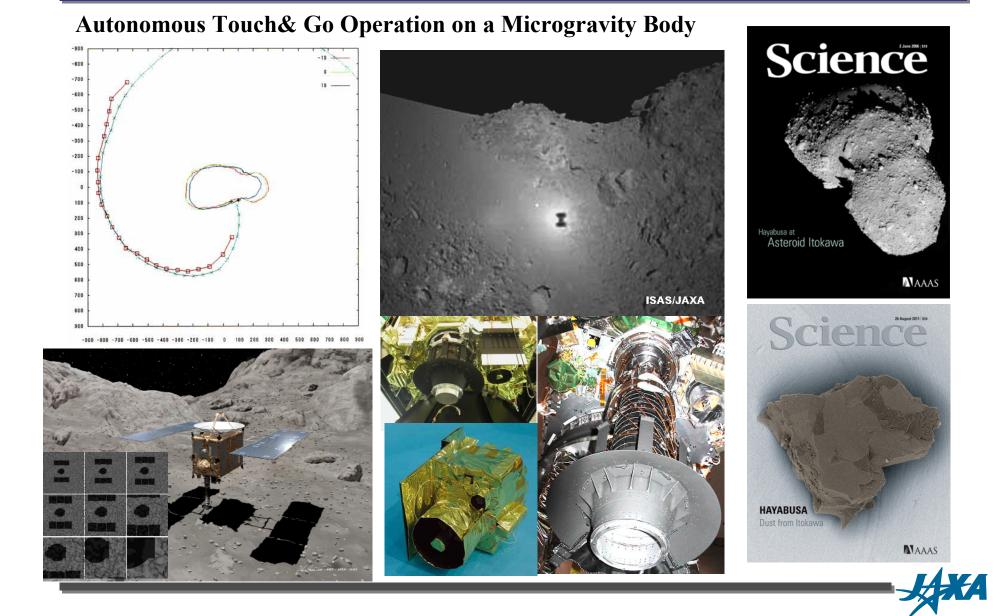
<Active partnership with international sample return missions by providing a package of key enabling technologies, rather than ad hoc instruments only>

- (1) Touchdown Sampling System
- (2) Flyby Sampling System and Capture Media Development
- (3) Ground Development and Calibration Experiment Facilities
- (4) Containment System, Contamination Control and Sample Preservation
- (5) Direct Earth Re-entry Capsule System
- (6) Ground Operation of Returned Capsule Recovery
- (7) Initial Sample Analysis and Curation Facility
- (8) Biosafety Sample Handling Operation and Facility by JAMSTEC Partner

Mission Heritage	(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)
Hayabusa-1 & 2	0		0	0	0	0	0	
Tanpopo		0	0				0	
Solar Power Sail	0			0	0	0	0	
JAMSTEC				0			0	0

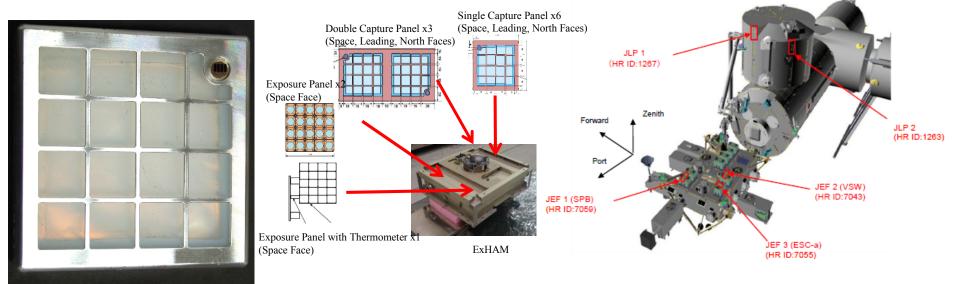


Key Technology-1: Touchdown Sampling System

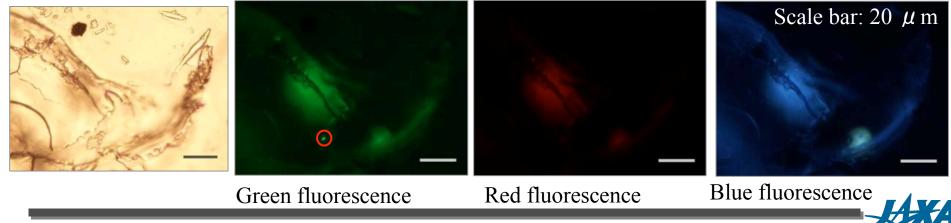


Key Technology-2: Flyby Sampling System and Capture Media Development

Japanese Double-Layered 0.01g/cc Aerogels in Tanpopo Capture Panel



Captured microbe colony embedded in clay minerals, which were shot into the tanpopo aerogel at 6km/s: Bleaching of green fluorescence from stained *D. radiodurans* R1 is faster than fluorescence from the glass



Key Technology-3: Ground Development and Calibration Experiment Facilities

Aerogel Surface

Ice Grain

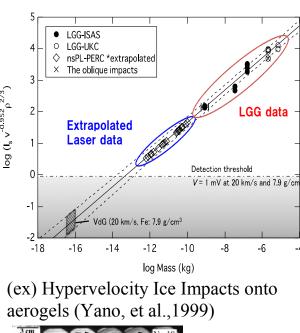
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Hypervelocity Impact Facilities



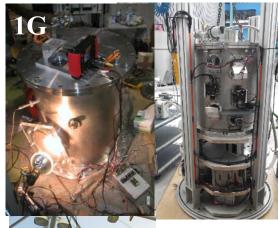
3.75MV Van de Graaf at Univ. Tokyo





Glass Bead

Microgravity Vacuum Chambers

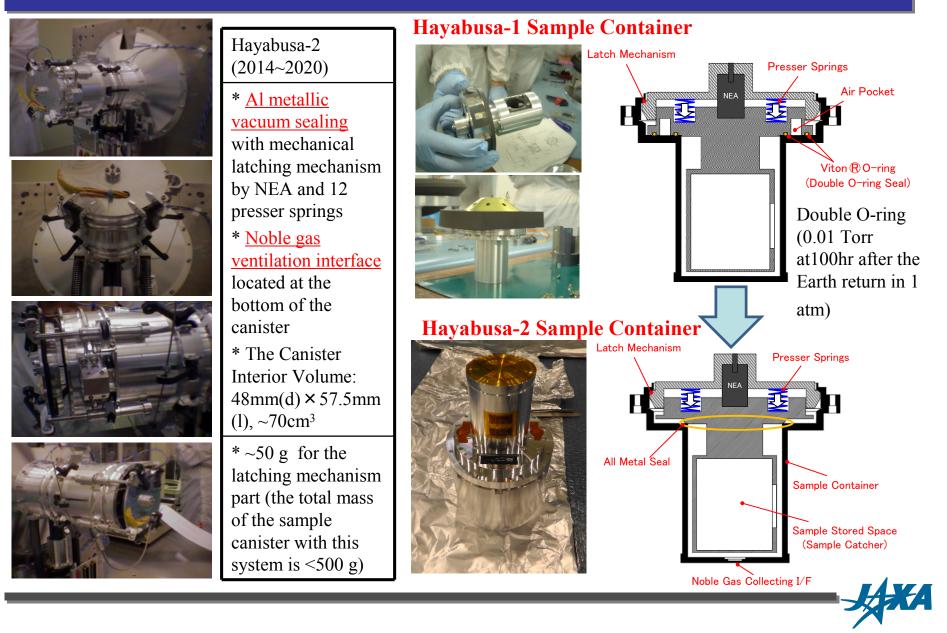




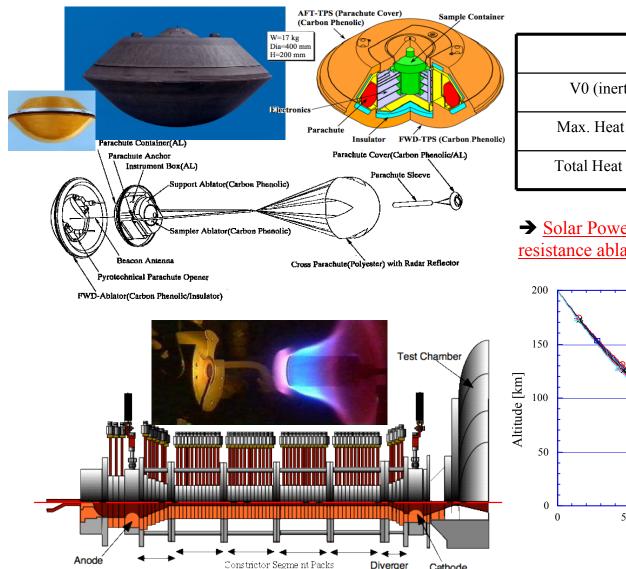




Key Technology-4: Containment System, Contamination Control and Sample Preservation



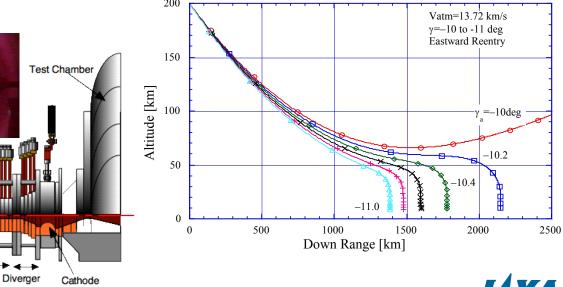
Key Technology-5: Direct Earth Re-entry Capsule System



Converger

	Hayabusa	Hayabusa-2
V0 (inertia)	12.06 km/s	11.88 km/s
Max. Heat Flux	15.7 MW/m ²	15.8 MW/m ²
Total Heat Input	256.7 MJ/m ²	235 MJ/m ²

→ <u>Solar Power Sail ERC needs higher heat</u> resistance ablator materials than the Hayabusa series



Key Technology-6: Ground Operation of Returned Capsule Recovery



Contingency Sterilization



Soil Sampling



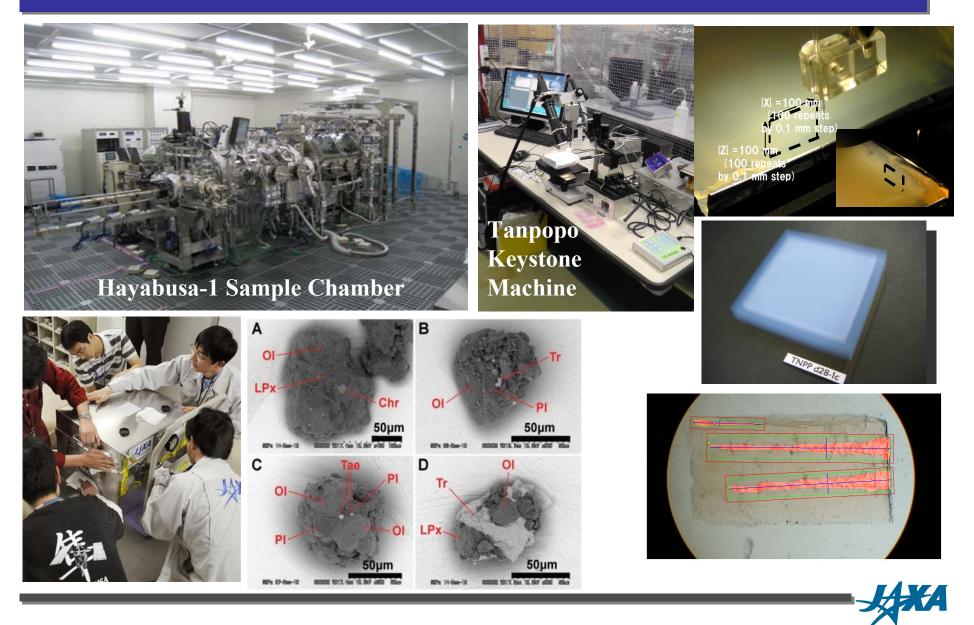
International Witness



Pre-Open XCT Inspection



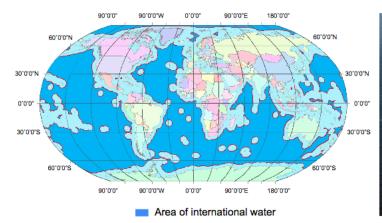
Key Technology-7: Initial Sample Analysis and Curation Facility



Key Technology-8: Biosafety Sample Handling Operation and Facility by JAMSTEC Partners

Subsurface hydrothermal microorganism sample return, contamination control and initial analysis on international waters of the Earth have been routinely and successfully conducted by the Research Vessel "Chikyu" of JAMSTEC Japan, as a part of the International Ocean Drilling Project (IODP) framework.

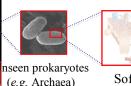
(A) International waters



(B) Onboard laboratory



BSL4 laboratory installed on the "Chikyu" to be provided by JAMSTEC partners





ea) Soft tissues (e.g. membrane biosignatures)

Present Routine Capability of the "Chikyu" Vessel



Helicopter Operation



Onboard Online Meeting Room



X-ray CT Scanner Prior to Opening



Anoxic Microbial Treatments



Microbial Culture Chamber

Stakeholders among Japanese Science and Engineering Community for New Sample Returns including Icy Plumes



Summary and Hopes for NASA

• JAXA/ISAS has a three decade history of deep space exploration and has built unique expertise for sample returns

• Keywords in future Japanese exploration vision include: round trip, outer planet region, astrobiology and international collaboration, which all are associate3d with planetary protection activities in different levels.

• The new JAXA planetary protection review board and the ISAS international collaboration sample return WG have been established in 2013. Both astrobiology community and space exploration stakeholders in Japan are rapidly growing, too.

• Key enabling technologies that the new WG can provide as a package range from sampling system, earth return capsule, to curation and biosafety treatment facilities in collaboration with the marine microbiology community led by JAMSTEC scientists.

→ The new WG hopes to have active collaboration with NASA through a comparable counterpart team with scientific background of astrobiology, engineering capability in deep space exploration and strong motivation to solve new planetary protection challenges such as backward contamination issues of icy plume returned samples from outer planet regions.

